

# Chapter 2 The One-Dimensional Scalar Wave Equation

2.1 - examine simplest differential equation describing wave propagation ↷

$$\frac{\partial^2 u}{\partial t^2} = c^2 \frac{\partial^2 u}{\partial x^2} \quad (2.1)$$

has as general solutions functions of form

$$u(x,t) = F(x+ct) + G(x-ct) \quad (2.2)$$

Take  $\frac{\partial}{\partial t}$  partial

↑ some functions that satisfy (2.1) ↑

$$\frac{\partial u}{\partial t} = \frac{dF(x+ct)}{d(x+ct)} \frac{\partial(x+ct)}{\partial t} + \frac{dG(x-ct)}{d(x-ct)} \frac{\partial(x-ct)}{\partial t}$$

↑ call this  $F'(x+ct)$       ↓  $c$       ↑ call this  $G'(x-ct)$       ↓  $-c$

$$= c F'(x+ct) - c G'(x-ct) \quad (2.3a)$$

Repeat process

$$\frac{\partial^2 u}{\partial t^2} = c^2 F''(x+ct) + c^2 G''(x-ct) \quad (2.3b)$$


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2.2 cont,

$$\frac{\partial u}{\partial x} = \frac{dF(x+ct)}{d(x+ct)} \frac{\partial(x+ct)}{\partial x} + \frac{dG(x-ct)}{d(x-ct)} \frac{\partial(x-ct)}{\partial x}$$

$\downarrow$   $\downarrow$   
 $\rightarrow 1$   $\rightarrow 1$

$$= F'(x+ct) + G'(x-ct) \tag{2.4a}$$

Repeat process

$$\frac{\partial^2 u}{\partial x^2} = F''(x+ct) + G''(x-ct) \tag{2.4b}$$

Now substitute back into (2.1)

$$\frac{\partial^2 u}{\partial t^2} = c^2 \frac{\partial^2 u}{\partial x^2}$$

$$c^2 F''(x+ct) + c^2 G''(x-ct) = c^2 [F''(x+ct) + G''(x-ct)] \tag{2.5}$$

↳ Proves that (2.2) satisfies (2.1)

2.3 Dispersion Relation

→ how does wavelength  $\lambda$  change w/ frequency  $f$ ?

Start by examining one solution to the 1D wave eqn (2.1)

$$u(x,t) = e^{j(\omega t - kx)} \leftarrow \text{fund traveling soln } G(x-ct) \tag{2.6}$$

### 2.3 cont.

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where the wavenumber  $k = \frac{2\pi}{\lambda} = \frac{\omega}{v_p}$  and  
the angular frequency  $\omega = 2\pi f$

$$\frac{du}{dt} = j\omega e^{j(\omega t - kx)}$$

$$\frac{du}{dx} = -jk e^{j(\omega t - kx)}$$

$$\frac{d^2u}{dt^2} = (j\omega)^2 e^{j(\omega t - kx)}$$

$$\frac{d^2u}{dx^2} = (-jk)^2 e^{j(\omega t - kx)}$$

Sub. into  
(2.1)

$$(j\omega)^2 e^{j(\omega t - kx)} = c^2 (-jk)^2 e^{j(\omega t - kx)}$$

divide out  
common terms

$$\omega^2 = c^2 k^2$$

$$k = \pm \frac{\omega}{c}$$

dispersion relation  
(2.7b)

OR

$$v_p = \frac{\omega}{k} = \pm c$$

(2.8)

phase  
velocity

\* Since  $v_p$  is constant wrt frequency,  
this solution is dispersionless.

2.3 cont.

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\* group velocity  $v_g = \frac{d\omega}{dk}$

\* In this case, start w/  $\omega^2 = c^2 k^2$  to get  $v_g$

$$\frac{d(\omega^2)}{dk} = \frac{d(c^2 k^2)}{dk}$$

$$2\omega \frac{d\omega}{dk} = c^2 (2k)$$

$$v_g = \frac{d\omega}{dk} = \frac{c^2 (2k)}{2\omega} = \frac{c^2 (\pm \omega/c)}{\omega}$$

$$\underline{v_g = \pm c} \quad (2.9)$$

↪ group velocity is also a constant (no frequency dependence)

# 2.4 Finite Differences

Do a Taylor's series expansion of  $u(x,t)$

Def'n of Taylor's series expansion of  $f(x)$  about point  $x=a$

$$f(x) = f(a) + (x-a) \left. \frac{\partial f(x)}{\partial x} \right|_{x=a} + \frac{(x-a)^2}{2!} \left. \frac{\partial^2 f(x)}{\partial x^2} \right|_{x=a}$$

$$+ \dots + \frac{(x-a)^{n-1}}{(n-1)!} \left. \frac{\partial^{n-1} f(x)}{\partial x^{n-1}} \right|_{x=a} + R_n \text{ or Remainder term}$$

Where

$$R_n = \frac{(x-a)^n}{n!} \left. \frac{\partial^n f(x)}{\partial x^n} \right|_{x=\xi} \quad \text{(Lagrange's form)}$$

lowercase  $\xi$

For our Taylor's series expansion of  $u(x,t)$ , we'll select a fixed time point  $t_n$  and expand about spatial point  $x_i$  to evaluate  $u(x,t)$  at  $x_i + \Delta x$  yielding

$$u(x_i + \Delta x, t_n) = u(x_i, t_n) + \frac{(\Delta x)^1}{1!} \left. \frac{\partial u}{\partial x} \right|_{x_i, t_n} + \frac{(\Delta x)^2}{2!} \left. \frac{\partial^2 u}{\partial x^2} \right|_{x_i, t_n}$$

$$+ \frac{(\Delta x)^3}{3!} \left. \frac{\partial^3 u}{\partial x^3} \right|_{x_i, t_n} + \frac{(\Delta x)^4}{4!} \left. \frac{\partial^4 u}{\partial x^4} \right|_{x_i, t_n}$$

$$u(x_i + \Delta x, t_n) = u(x_i, t_n) + \Delta x \left. \frac{\partial u}{\partial x} \right|_{x_i, t_n} + \frac{\Delta x^2}{2} \left. \frac{\partial^2 u}{\partial x^2} \right|_{x_i, t_n}$$

$$+ \frac{\Delta x^3}{6} \left. \frac{\partial^3 u}{\partial x^3} \right|_{x_i, t_n} + \frac{\Delta x^4}{24} \left. \frac{\partial^4 u}{\partial x^4} \right|_{x_i, t_n} \quad (2.10a)$$

↑  
error/remainder term

2.4 cont.

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Repeat process for  $(x_i - \Delta x, t_n)$  to get

$$\begin{aligned} u(x_i - \Delta x, t_n) &= u(x_i, t_n) + \frac{(x_i - \Delta x - x_i)^{\overbrace{-\Delta x}^{\uparrow}}}{1!} \frac{\partial u}{\partial x} \Big|_{x_i, t_n} + \frac{(-\Delta x)^2}{2!} \frac{\partial^2 u}{\partial x^2} \Big|_{x_i, t_n} \\ &\quad + \frac{(-\Delta x)^3}{3!} \frac{\partial^3 u}{\partial x^3} \Big|_{x_i, t_n} + \frac{(-\Delta x)^4}{4!} \frac{\partial^4 u}{\partial x^4} \Big|_{\mathcal{F}_2, t_n} \\ &= u(x_i, t_n) - \Delta x \frac{\partial u}{\partial x} \Big|_{x_i, t_n} + \frac{\Delta x^2}{2} \frac{\partial^2 u}{\partial x^2} \Big|_{x_i, t_n} \\ &\quad - \frac{\Delta x^3}{6} \frac{\partial^3 u}{\partial x^3} \Big|_{x_i, t_n} + \frac{\Delta x^4}{24} \frac{\partial^4 u}{\partial x^4} \Big|_{\mathcal{F}_2, t_n} \quad (2.10b) \end{aligned}$$

Now add (2.10a) to (2.10b)

$$x_i - \Delta x < \mathcal{F}_2 < x_i$$

$$\begin{aligned} u(x_i + \Delta x, t_n) + u(x_i - \Delta x, t_n) &= 2u(x_i, t_n) + 0 + \Delta x^2 \frac{\partial^2 u}{\partial x^2} \Big|_{x_i, t_n} \\ &\quad + 0 + \underbrace{\frac{\Delta x^4}{24} \left( \frac{\partial^4 u}{\partial x^4} \Big|_{\mathcal{F}_1, t_n} + \frac{\partial^4 u}{\partial x^4} \Big|_{\mathcal{F}_2, t_n} \right)}_{\text{combine}} \end{aligned}$$

$$(2.11) \quad u(x_i + \Delta x, t) + u(x_i - \Delta x, t) = 2u(x_i, t_n) + \Delta x^2 \frac{\partial^2 u}{\partial x^2} \Big|_{x_i, t_n} + \frac{\Delta x^4}{12} \frac{\partial^4 u}{\partial x^4} \Big|_{\mathcal{F}_2, t_n}$$

$$\text{where } x_i - \Delta x < \mathcal{F}_3 < x_i + \Delta x$$

Now, solve (2.11) for  $\frac{\partial^2 u}{\partial x^2} \Big|_{x_i, t_n}$

## 2.4 cont.

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$$(2.12) \quad \left. \frac{\partial^2 U}{\partial x^2} \right|_{x_i, t_n} = \frac{U(x_i + \Delta x, t_n) - 2U(x_i, t_n) + U(x_i - \Delta x, t_n)}{\Delta x^2} - \underbrace{\frac{\Delta x^4}{12\Delta x^2} \left. \frac{\partial^4 U}{\partial x^4} \right|_{x_i, t_n}}_{\text{Remainder term}} \rightarrow O[(\Delta x)^2]$$

Using the standard FDTD notation where  $x_i = i\Delta x$  and  $t_n = n\Delta t$  and  $U(x_i, t_n) = U_i^n$

$$\left. \frac{\partial^2 U}{\partial x^2} \right|_{x_i, t_n} = \frac{U_{i+1}^n - 2U_i^n + U_{i-1}^n}{\Delta x^2} + O[(\Delta x)^2] \quad (2.13)$$

Similarly, we can do a Taylor's series expansion of  $u(x, t)$  for a fixed spatial point  $x_i$  and expand about the time point  $t_n$  both forward ( $t_n + \Delta t$ ) and backward ( $t_n - \Delta t$ ) in time to get

$$\left. \frac{\partial^2 U}{\partial t^2} \right|_{x_i, t_n} = \frac{U_i^{n+1} - 2U_i^n + U_i^{n-1}}{\Delta t^2} + O[(\Delta t)^2] \quad (2.14)$$

\* (2.13) and (2.14) are second-order accurate ( $\Delta x^2, \Delta t^2$ ) central-difference approximations to  $\frac{\partial^2 U}{\partial x^2}$  and  $\frac{\partial^2 U}{\partial t^2}$ .

## 2.5 Finite-Difference Approximation of the Scalar Wave Equation

Now, substitute (2.13) and (2.14) into the 1D wave equation (2.1) to get

$$\frac{u_i^{n+1} - 2u_i^n + u_i^{n-1}}{\Delta t^2} + O[(\Delta t)^2] = c^2 \left\{ \frac{u_{i+1}^n - 2u_i^n + u_{i-1}^n}{\Delta x^2} + O[(\Delta x)^2] \right\}$$

Dropping the error/remainder terms and solving for  $u_i^{n+1}$  (step forward in time) yields

$$u_i^{n+1} = \underbrace{\left( \frac{c\Delta t}{\Delta x} \right)^2 [u_{i+1}^n - 2u_i^n + u_{i-1}^n]}_{\text{all these terms are known at time } t_n} + 2u_i^n - u_i^{n-1} \quad (2.16)$$

↶ calculate future values using information from current + past time steps

If we choose our time interval  $\Delta t$  such that  $c\Delta t = \Delta x$  (magic time-step)

$$u_i^{n+1} = u_{i+1}^n + u_{i-1}^n - u_i^{n-1} \quad (2.17)$$

↶ Not approximation anymore

Proof

Going back to  $u(x,t) = F(x+ct) + G(x-ct)$ , and evaluating at  $(x_i, t_n)$  gives

$$U_i^n = F(x_i + ct_n) + G(x_i - ct_n).$$

Next, substitute this into (2.17) to get

$$\left[ \begin{array}{l} F(x_i + ct_{n+1}) \\ + G(x_i - ct_{n+1}) \end{array} \right] = \left[ \begin{array}{l} F(x_{i+1} + ct_n) \\ + G(x_{i+1} - ct_n) \end{array} \right] + \left[ \begin{array}{l} F(x_{i-1} + ct_n) \\ + G(x_{i-1} - ct_n) \end{array} \right] - \left[ \begin{array}{l} F(x_i + ct_{n-1}) \\ + G(x_i - ct_{n-1}) \end{array} \right] \quad (2.18a)$$

$$RHS = \left\{ \begin{array}{l} F[(i+1)\Delta x + c\Delta t] \\ + G[(i+1)\Delta x - c\Delta t] \end{array} \right\} + \left\{ \begin{array}{l} F[(i-1)\Delta x + c\Delta t] \\ + G[(i-1)\Delta x - c\Delta t] \end{array} \right\} - \left\{ \begin{array}{l} F[i\Delta x + c(n-1)\Delta t] \\ + G[i\Delta x - c(n-1)\Delta t] \end{array} \right\} \quad (2.18b)$$

Substitute in the magic time-step  $c\Delta t = \Delta x$ , to get

$$RHS = \left\{ \begin{array}{l} F[(i+1)\Delta x + n\Delta x] \\ + G[(i+1)\Delta x - n\Delta x] \end{array} \right\} + \left\{ \begin{array}{l} F[(i-1)\Delta x + n\Delta x] \\ + G[(i-1)\Delta x - n\Delta x] \end{array} \right\} - \left\{ \begin{array}{l} F[i\Delta x + (n-1)\Delta x] \\ + G[i\Delta x - (n-1)\Delta x] \end{array} \right\} \quad (2.18c)$$

cancel

$$RHS = F[(i+1+n)\Delta x] + G[(i-1-n)\Delta x] \quad (2.18d)$$

↑ original term moved to right by  $\Delta x$       ↑ original term moved left  $\Delta x$

How can we prove this?

$$\begin{aligned}u_i^{n+1} &= F(x_i + ct_{n+1}) + G(x_i - ct_{n+1}) \\ &= F[i\Delta x + c(n+1)\Delta t] + G[i\Delta x - c(n+1)\Delta t] \quad (2.19a)\end{aligned}$$

Again, substitute in the magic time-step  
( $c\Delta t = \Delta x$ )

$$\underline{u_i^{n+1} = F[(i+n+1)\Delta x] + G[(i-n-1)\Delta x]} \quad (2.19b)$$

same as (2.18d)  $\nearrow$

$\Rightarrow$  Approximation errors/remainders  
cancel out at magic time-step  $\circ$

## 2.6 Numerical Dispersion Relation

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→ want to characterize dispersion due to the finite-difference solution/approximation to the 1D scalar wave equation

Again, consider  $u(x,t) = e^{j(\omega t - kx)}$  (2.6)

as represented numerically w/  $k$  replaced with  $\tilde{k} = \tilde{k}_{\text{real}} + j\tilde{k}_{\text{imag}}$  (numerical wave number).

This gives

$$u_i^n = e^{j(\omega n \Delta t - \tilde{k} i \Delta x)} = e^{\tilde{k}_{\text{imag}} i \Delta x} e^{j(\omega n \Delta t - \tilde{k}_{\text{real}} i \Delta x)} \quad (2.20)$$

Ideally,  $\tilde{k}_{\text{imag}} = 0$  and  $\tilde{k}_{\text{real}} = k$ .

Next, substitute (2.20) into the finite-difference approximation of (2.16) to get

$$e^{j[\omega(n+1)\Delta t - \tilde{k} i \Delta x]} = \left(\frac{c\Delta t}{\Delta x}\right)^2 \left\{ e^{j[\omega n \Delta t - \tilde{k}(i+1)\Delta x]} - 2e^{j[\omega n \Delta t - \tilde{k} i \Delta x]} + e^{j[\omega n \Delta t - \tilde{k}(i-1)\Delta x]} \right\} + 2e^{j[\omega n \Delta t - \tilde{k} i \Delta x]} - e^{j[\omega(n-1)\Delta t - \tilde{k} i \Delta x]} \quad (2.21a)$$

divide out common  $e^{j[\omega n \Delta t - \tilde{k} i \Delta x]}$  terms, to get

## 2.6 conts

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$$e^{j\omega\Delta t} = \left(\frac{c\Delta t}{\Delta x}\right)^2 \left[ e^{-j\tilde{k}\Delta x} - 2 + e^{j\tilde{k}\Delta x} \right] + 2 - e^{-j\omega\Delta t} \quad (2.21b)$$

Do a little algebra to allow Euler's ID for cosine terms (i.e.,  $\cos A = \frac{e^{jA} + e^{-jA}}{2}$ ) to be applied

$$\frac{e^{j\omega\Delta t} + e^{-j\omega\Delta t}}{2} = \left(\frac{c\Delta t}{\Delta x}\right)^2 \left[ \frac{e^{j\tilde{k}\Delta x} + e^{-j\tilde{k}\Delta x}}{2} - 1 \right] + 1 \quad (2.21c)$$

This gives

$$\cos(\omega\Delta t) = \left(\frac{c\Delta t}{\Delta x}\right)^2 \left[ \cos(\tilde{k}\Delta x) - 1 \right] + 1 \quad (2.22)$$

and

$$\tilde{k} = \frac{1}{\Delta x} \cos^{-1} \left\{ 1 + \left(\frac{\Delta x}{c\Delta t}\right)^2 [\cos(\omega\Delta t) - 1] \right\} \quad (2.23)$$

Now we can find how choices of  $\Delta x$  and  $\Delta t$  will cause  $\tilde{k}$  to be different than  $k = \pm \frac{\omega}{c}$  (2.7b)

## 2.6.1 Case 1: Very Fine Sampling in Time + Space 13

→ assume  $\Delta t$  +  $\Delta x$  are very small wrt  $T$  and  $\lambda$  respectively

→ can truncate the Taylor's Series expansion of the cosine function to two terms, e.g.

$$\cos x = 1 - \frac{x^2}{2!} + \underbrace{\frac{x^4}{4!} - \frac{x^6}{6!} + \dots}_{\text{neglect}}$$

So (2.23) becomes

$$\tilde{k} = \frac{1}{\Delta x} \cos^{-1} \left\{ 1 + \left( \frac{\Delta x}{c \Delta t} \right)^2 \left[ \cos \left( \sqrt{\frac{(\omega \Delta t)^2}{2}} \right) - 1 \right] \right\}$$

$$\approx \frac{1}{\Delta x} \cos^{-1} \left\{ 1 + \left( \frac{\Delta x}{c \Delta t} \right)^2 \left( - \frac{(\omega \Delta t)^2}{2} \right) \right\}$$

$$\approx \frac{1}{\Delta x} \cos^{-1} \left[ 1 - \underbrace{\left( \frac{\omega}{c} \right)^2}_{k^2} \frac{\Delta x^2}{2} \right]$$

$$\tilde{k} \approx \frac{1}{\Delta x} \cos^{-1} \left[ 1 - \frac{(k \Delta x)^2}{2} \right] \quad (2.24a)$$

Again use  $\cos x \approx 1 - \frac{x^2}{2}$  for  $|x| \ll 1$

$$\cos^{-1}(\cos x) = x \approx \cos^{-1} \left( 1 - \frac{x^2}{2} \right)$$

to get

$$\tilde{k} \approx \frac{1}{\Delta x} (k \Delta x) = k \quad (2.24b)$$

2.6.1 cont.

→ As might be expected, for very fine sampling, the numerical solution converges w/ the continuous solution (i.e., no dispersion!)

2.6.2 Case 2: Magic Time-Step

→ let  $\Delta x = c \Delta t$  in (2.23)

$$\tilde{k} = \frac{1}{\underset{\substack{\uparrow \\ c \Delta t}}{\Delta x}} \cos^{-1} \left\{ 1 + \left( \frac{\omega \Delta t}{c \Delta t} \right)^2 [\cos(\omega \Delta t) - 1] \right\} \quad (2.25a)$$

$$= \frac{1}{c \Delta t} \cos^{-1} \left\{ \cos(\omega \Delta t) \right\} = \frac{\omega \Delta t}{c \Delta t}$$

$\tilde{k} = \omega / c = k$  (2.25b)

→ Again, No Dispersion! However, there are no restrictions on  $\Delta x + \Delta t$  other than they obey  $\Delta x = c \Delta t$ .

### 2.6.3 Case 3: Dispersive Wave Propagation

→ What happens if  $\Delta x$  +  $\Delta t$  don't satisfy case 1 or 2?

ex. let  $\Delta x = \lambda_0/10$  and choose  $c\Delta t = \Delta x/2$

( $\frac{\Delta x}{c\Delta t} = 2 \rightarrow$  two time steps to travel  $\Delta x$ )

$$\tilde{k} = \frac{1}{\Delta x} \cos^{-1} \left\{ 1 + \underbrace{\left(\frac{\Delta x}{c\Delta t}\right)^2}_{2^2} \left[ \cos\left(\omega \Delta t \frac{\Delta x}{2c}\right) - 1 \right] \right\}$$

$$= \frac{1}{\Delta x} \cos^{-1} \left\{ 1 + 4 \left[ \cos\left(\frac{\omega \Delta x}{2c}\right) - 1 \right] \right\}$$

$$\uparrow k = \omega/c = \frac{2\pi}{\lambda_0}$$

$$= \frac{1}{\Delta x} \cos^{-1} \left\{ 1 + 4 \left[ \cos\left(\frac{2\pi}{\lambda_0} \frac{\Delta x}{2} \frac{\lambda_0}{10}\right) - 1 \right] \right\}$$

$$= \frac{1}{\Delta x} \cos^{-1} \left\{ 1 + 4 \left[ \cos\left(\frac{\pi}{10}\right) - 1 \right] \right\}$$

$$= \frac{1}{\Delta x} \cos^{-1} \{ 0.804226 \} = \frac{0.636424}{\Delta x} \quad (2.265)$$

$\Delta x \rightarrow \lambda_0/10$

$$\tilde{k} = \frac{6.364242196}{\lambda_0}$$

1.29% larger!  
than  $k$

$$k = \frac{2\pi}{\lambda_0} = \frac{6.2831853}{\lambda_0}$$

2.6.3 cont,

ex, cont.

looking at the numerical phase velocity  $\tilde{v}_p$

$$\tilde{v}_p = \frac{\omega}{\tilde{k}} = \frac{2\pi f \uparrow \frac{c}{\lambda_0}}{\left(\frac{6.304242196}{\lambda_0}\right)} = \underline{\underline{0.98726 c}} \quad (2.276)$$

↳ wave propagates slower than speed of light in the numerical grid (1.27% slower)

ex. let  $\Delta x = \frac{\lambda_0}{20}$  w/  $c\Delta t = \frac{\Delta x}{2}$  again

$$\tilde{k} = \frac{1}{\Delta x} \cos^{-1} \left\{ 1 + 4 \left[ \cos \left( \frac{2\pi}{\lambda_0} \frac{\Delta x}{2} \right) - 1 \right] \right\}$$

$\rightarrow \frac{\lambda_0}{20}$

$$= \frac{20}{\lambda_0} \cos^{-1} \{ 0.950753362 \}$$

$$\tilde{k} = \frac{6.302776}{\lambda_0} \quad \leftarrow \text{only } 0.312\% \text{ larger than}$$

$k = \frac{2\pi}{\lambda_0}$

$$\tilde{v}_p = \frac{2\pi \frac{c}{\lambda_0}}{\left(\frac{6.303}{\lambda_0}\right)} = \underline{\underline{0.99689 c}} \quad \leftarrow \text{much better match}$$

(0.31% slower)

2.6.3 cont.

Now, let's consider the general case. Start by defining a numerical stability factor or

$$\underline{\text{Courant number}} \equiv S = \frac{c\Delta t}{\Delta x} \quad (2.28a)$$

and the grid sampling resolution

$$\underline{N_\lambda} = \frac{\lambda_0}{\Delta x} \quad (2.28b)$$

With these the numerical wavenumber can be written

$$\tilde{k} = \frac{1}{\Delta x} \cos^{-1} \left\{ 1 + \left(\frac{1}{S}\right)^2 \left[ \cos\left(\frac{2\pi S}{N_\lambda}\right) - 1 \right] \right\}$$
$$= \frac{1}{\Delta x} \cos^{-1} \left\{ \frac{\varphi}{\zeta} \right\} = \frac{1}{\Delta x} \left[ \frac{\pi}{2} - \sin^{-1} \left( \frac{\varphi}{\zeta} \right) \right] \quad (2.29a)$$

↑ zeta

where

$$\varphi = 1 + \left(\frac{1}{S}\right)^2 \left[ \cos\left(\frac{2\pi S}{N_\lambda}\right) - 1 \right]$$

↪ If  $\varphi < -1$ ,  $\tilde{k}$  becomes complex

This translates to

$$N_\lambda \Big|_{\text{transition}} = \frac{2\pi S}{\cos^{-1}(1 - 2S^2)} \quad (2.30)$$

### 2.6.3 cont.

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If  $N_\lambda > N_\lambda|_{\text{trans}}$ ,  $\tilde{\kappa}$  is real. Conversely,

if  $N_\lambda < N_\lambda|_{\text{trans}}$ ,  $\tilde{\kappa}$  is imaginary.

ex. In the previous example,  $c\Delta t = \frac{\Delta x}{2}$ .

$$\text{Therefore, } S = \frac{c\Delta t}{\Delta x} = 0.5$$

$$N_\lambda|_{\text{trans}} = \frac{2\pi(0.5)}{\cos^{-1}(1-2(0.5)^2)} = 3$$

Let's choose  $N_\lambda = 2.5$  w/  $S = 0.5$ , and calculate  $\tilde{\kappa}$

$$\begin{aligned} \rho &= 1 + \left(\frac{1}{0.5}\right)^2 \left[ \cos\left(\frac{2\pi(0.5)}{2.5}\right) - 1 \right] \\ &= -1.763932 \end{aligned}$$

$$\tilde{\kappa} = \frac{1}{\Delta x} \cos^{-1}\{-1.763932\} = \frac{\pi - j1.16845}{\Delta x \rightarrow \lambda_0/2.5}$$

$$\tilde{\kappa} = \frac{2.5\pi - j2.921}{\lambda_0} \quad \leftarrow \text{Nowhere close to } \kappa = \frac{2\pi}{\lambda_0}$$

Real  $\tilde{K}$  ( $N_1 > N_1|_{\text{transition}}$ )

$$\tilde{K} = \tilde{K}_{\text{real}} = \frac{1}{\Delta x} \cos^{-1} \left\{ 1 + \left(\frac{1}{5}\right)^2 \left[ \cos\left(\frac{2\pi 5}{N_1}\right) - 1 \right] \right\}$$

$$\tilde{K}_{\text{imag}} = 0 \quad \leftarrow \text{No numerical wave amplification or attenuation } (e^{\tilde{K}_{\text{imag}} \Delta x} = e^0 = 1)$$

$\tilde{K}$  complex ( $N_1 < N_1|_{\text{transition}}$ )

$$\begin{aligned} \tilde{K}_{\text{real}} &= \frac{\pi}{\Delta x} \\ \tilde{K}_{\text{imag}} &= \frac{1}{\Delta x} \ln(-\beta - \sqrt{\beta^2 - 1}) \end{aligned} \quad (2.36)$$

$$\tilde{V}_p = \frac{\omega}{\tilde{K}_{\text{real}}} = \frac{\omega}{(\pi/\Delta x)} = \frac{2f\lambda_0}{N_1} = \frac{2}{N_1} < \quad (2.37a)$$

and wave-amplitude multiplier per  $\Delta x$

$$e^{\tilde{K}_{\text{imag}} \Delta x} \equiv e^{-\alpha \Delta x} = e^{\ln(-\beta - \sqrt{\beta^2 - 1})} = \underline{\underline{-\beta - \sqrt{\beta^2 - 1}}} \quad (2.37b)$$

Note: If  $N_1 < 2$ ,  $\tilde{V}_p > 1$ !

(Remember Nyquist Sampling Rate  $f_{\text{max}} = \frac{1}{2\Delta t} = \frac{f_{\text{sample}}}{2}$ )

### 2.6.3 cont.

20

Applying this observation

$$d_{0,\min} = \frac{c}{f_{\max}} = 2c \Delta t$$

which leads to

$$\underline{N_{\lambda,\min} = \frac{d_{0,\min}}{\Delta x} = \frac{2c \Delta t}{\Delta x} = 2S \quad (2.38b)}$$

For stability  $S \leq 1$ , this leads to

$$N_{\lambda,\min} = 2S \leq 2 \quad (2.38c)$$

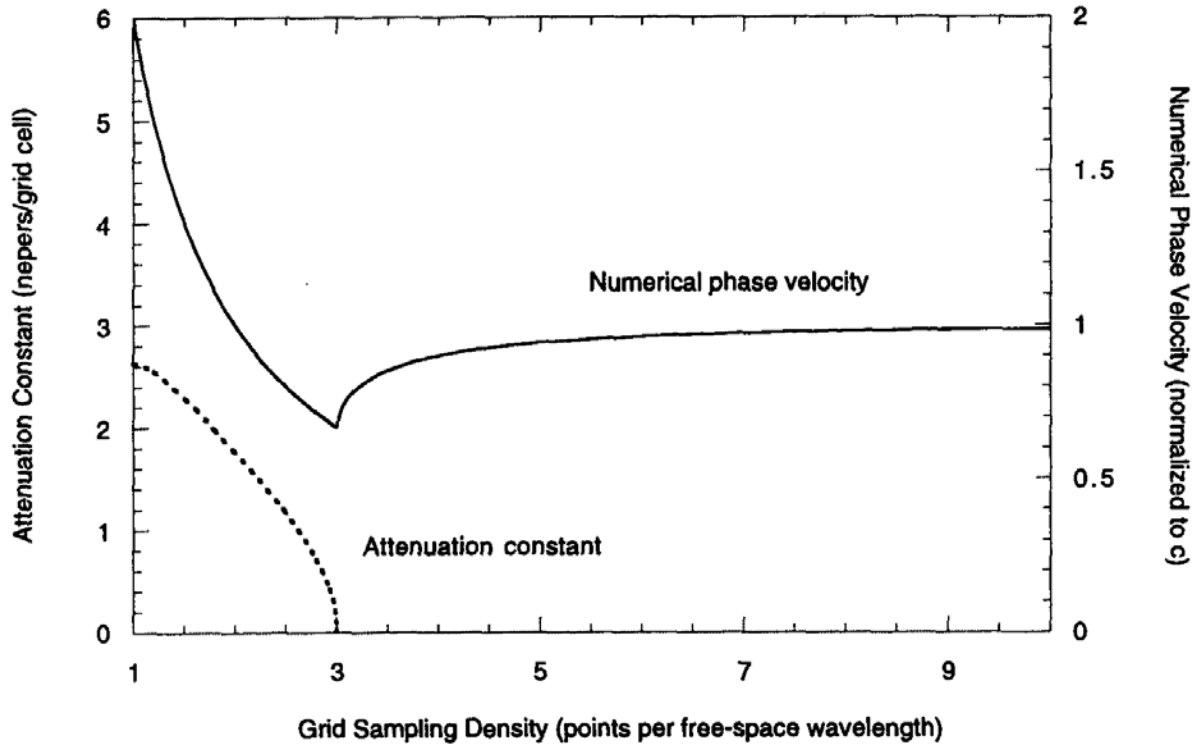
and

$$\tilde{v}_{p,\max} = \frac{2}{N_{\lambda,\min}} c = \frac{1}{S} c \geq c \quad (2.39a)$$

$$\underline{\tilde{v}_{p,\max} = \frac{\Delta x}{\Delta t}} \quad (2.39b)$$

\* Tells us wave can go no more than  $1 \Delta x$  per time increment  $\Delta t$ .

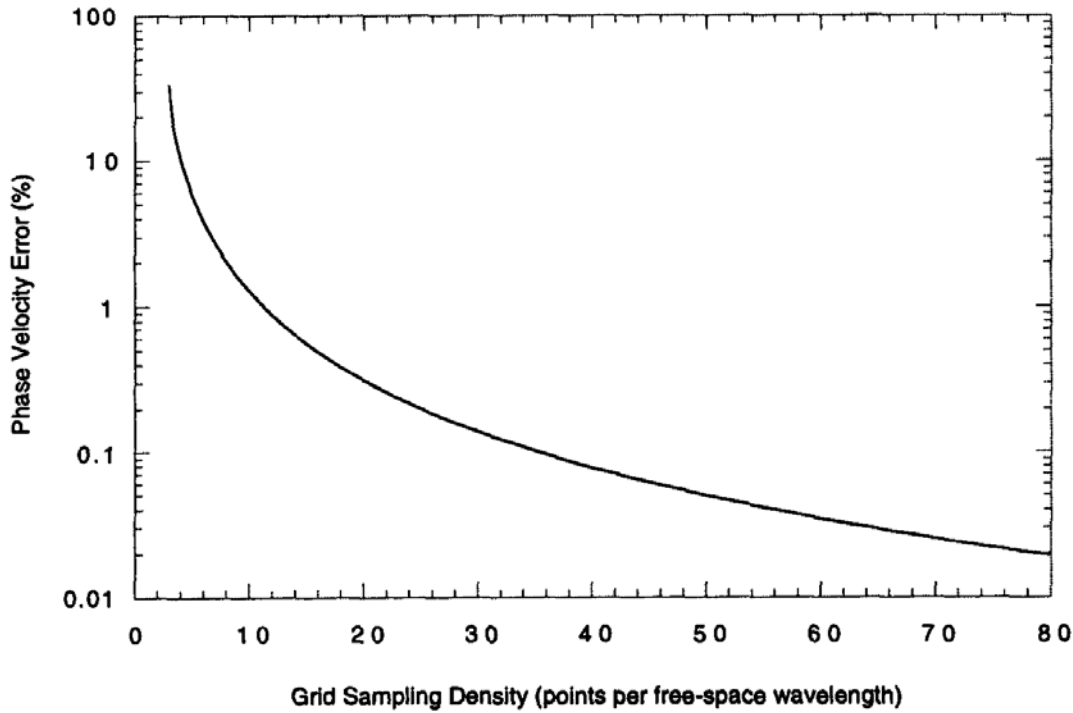
\* Also,  $\tilde{v}_{p,\max}$  doesn't depend on material properties



**Fig. 2.1** Variation of the normalized numerical phase velocity  $\bar{v}_p/c$  and attenuation per grid cell  $\alpha\Delta x$  as a function of the grid sampling density ( $1 \leq N_\lambda \leq 10$ ) for a Courant stability factor  $S = 0.5$ .

### Comments:

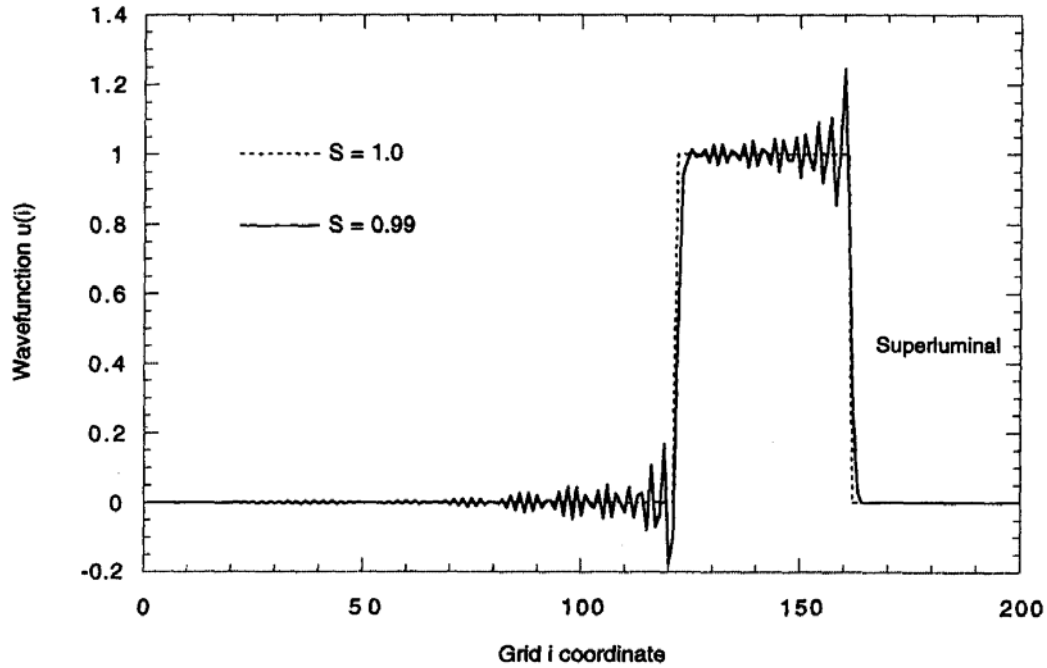
- In this case,  $N_\lambda \big|_{\text{transition}} = 3$ .
- Note how the attenuation constant disappears for  $N_\lambda > N_\lambda \big|_{\text{transition}} = 3$ .
- Note how the attenuation constant approaches 2.639 as  $N_\lambda \rightarrow 1$ .
- Note how the numerical phase velocity is greater than the speed of light for  $N_\lambda < 2$ .
- Note how the numerical phase velocity is less than the speed of light for  $N_\lambda < 2$ .
- Note how the numerical phase velocity approaches the speed of light as  $N_\lambda \rightarrow \infty$ .



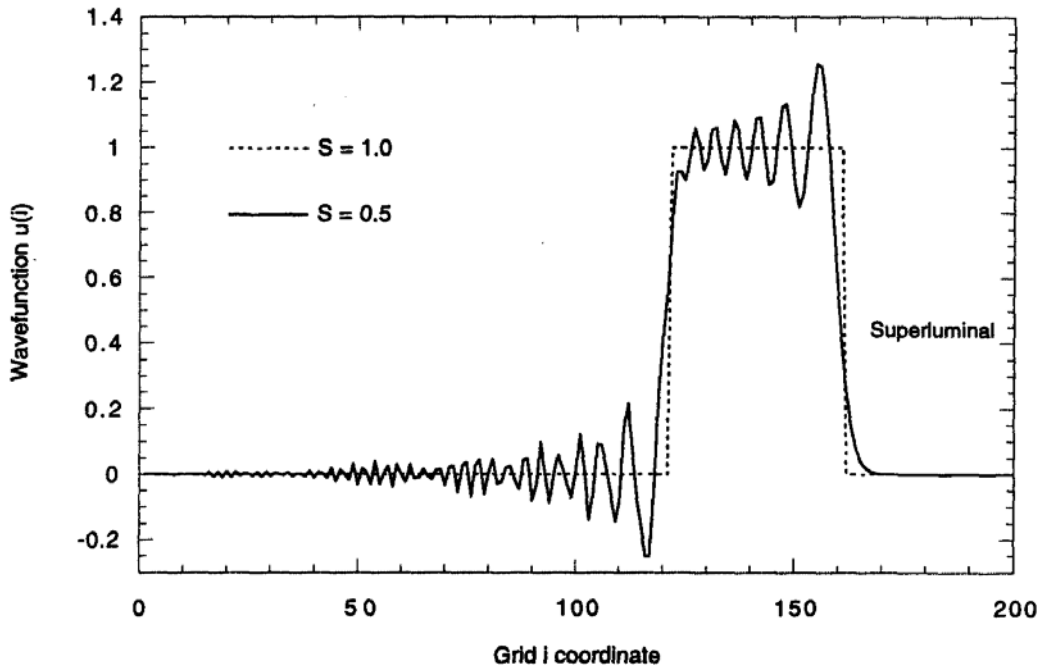
**Fig. 2.2** Percent numerical phase-velocity error relative to the free-space speed of light as a function of the grid sampling density ( $3 \leq N_\lambda \leq 80$ ) for a Courant stability factor  $S = 0.5$ .

### Comments:

- Note how the % error between the numerical phase velocity and free-space phase velocity approaches 0 as  $N_\lambda \rightarrow \infty$ .
- Decrease is proportional to  $1/N_\lambda^2$  due to second-order accuracy of finite-difference approximation.



(a) Comparison of calculated pulse propagation for  $S = 1$  and  $S = 0.99$ .



(b) Comparison of calculated pulse propagation for  $S = 1$  and  $S = 0.5$ .

**Fig. 2.3** Effect of numerical dispersion upon a rectangular pulse propagating in free space for different Courant stability factors: (a)  $S = 1$  ( $\Delta t$  equal to the magic time-step) and  $S = 0.99$  ( $\Delta t$  just 1% below the magic time-step); (b)  $S = 1$  and  $S = 0.5$  ( $\Delta t$  50% below the magic time-step).

**Fig. 2.3a) comments:**

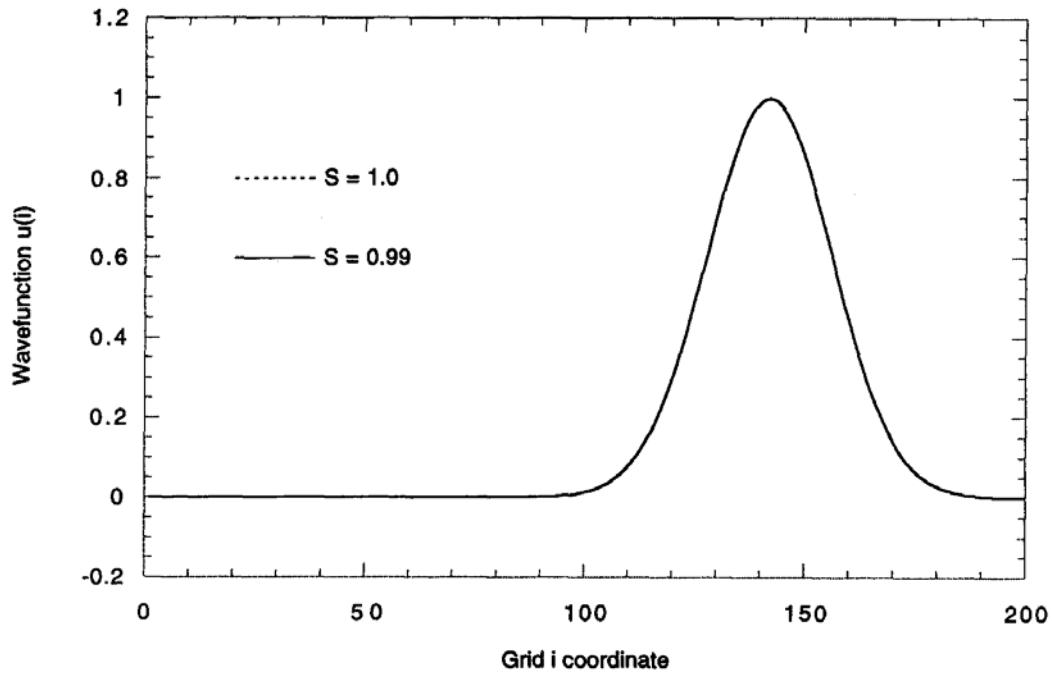
- In this case, the input signal is a rectangular pulse which is 40 spatial cells wide. From Fourier transform theory we know that this signal has infinite frequency content that does not damp out quickly, e.g.,

$$p_{\tau}(t) \leftrightarrow \tau \operatorname{sinc}\left(\frac{\tau\omega}{2\pi}\right).$$

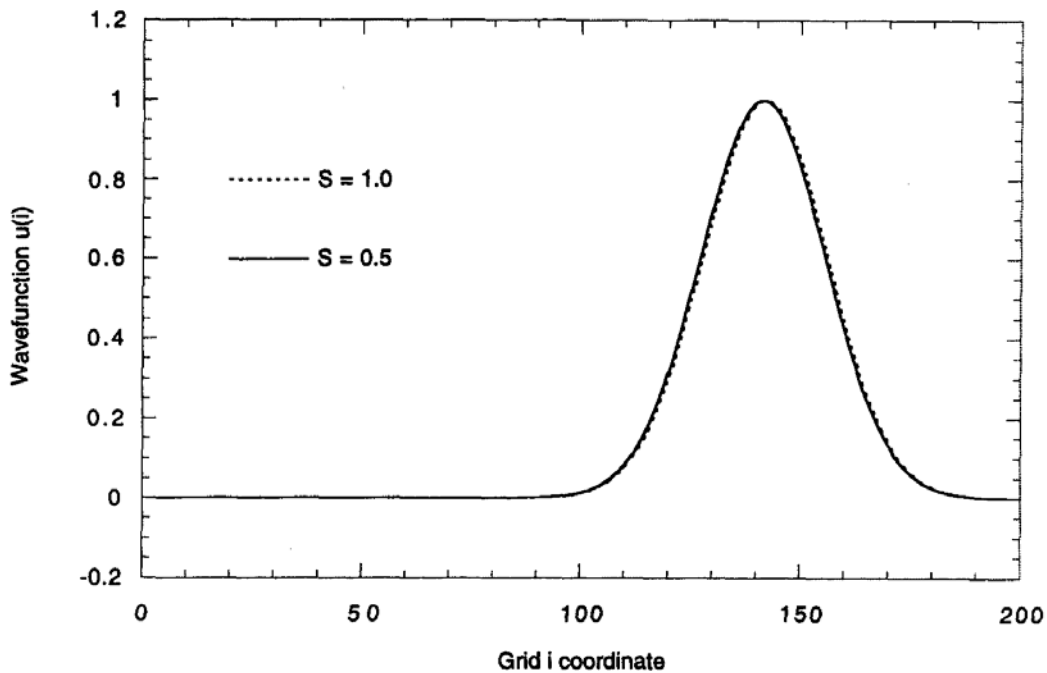
- For the magic time-step ( $S = 1 = c\Delta t / \Delta x$ ), the pulse is perfectly recreated since  $\tilde{v}_p \equiv c$  is constant at all frequencies (no dispersion).
- However, when  $S = 0.99$ ,  $\tilde{v}_p$  is NOT constant at all frequencies (dispersion) and we see delayed ringing at the edges of the pulse (high frequency content is not adequately sampled).
- Also, when  $S = 0.99$ , we see some superluminal signal content **before** the leading edge of the pulse (i.e.,  $\tilde{v}_p > c$ ).

**Fig. 2.3b) comments:**

- Here, we see what happens when  $S = 0.5$ .
- Much worse ringing with  $S = 0.5$ .
- More superluminal signal content with  $S = 0.5$ .



(a) Comparison of calculated pulse propagation for  $S = 1$  and  $S = 0.99$ .



(b) Comparison of calculated pulse propagation for  $S = 1$  and  $S = 0.5$ .

**Fig. 2.4** Effect of numerical dispersion upon a Gaussian pulse propagating in free space for different Courant stability factors: (a)  $S = 1$  ( $\Delta t$  equal to the magic time-step) and  $S = 0.99$  ( $\Delta t$  just 1% below the magic time-step); (b)  $S = 1$  and  $S = 0.5$  ( $\Delta t$  50% below the magic time-step).

**Fig. 2.4a) comments:**

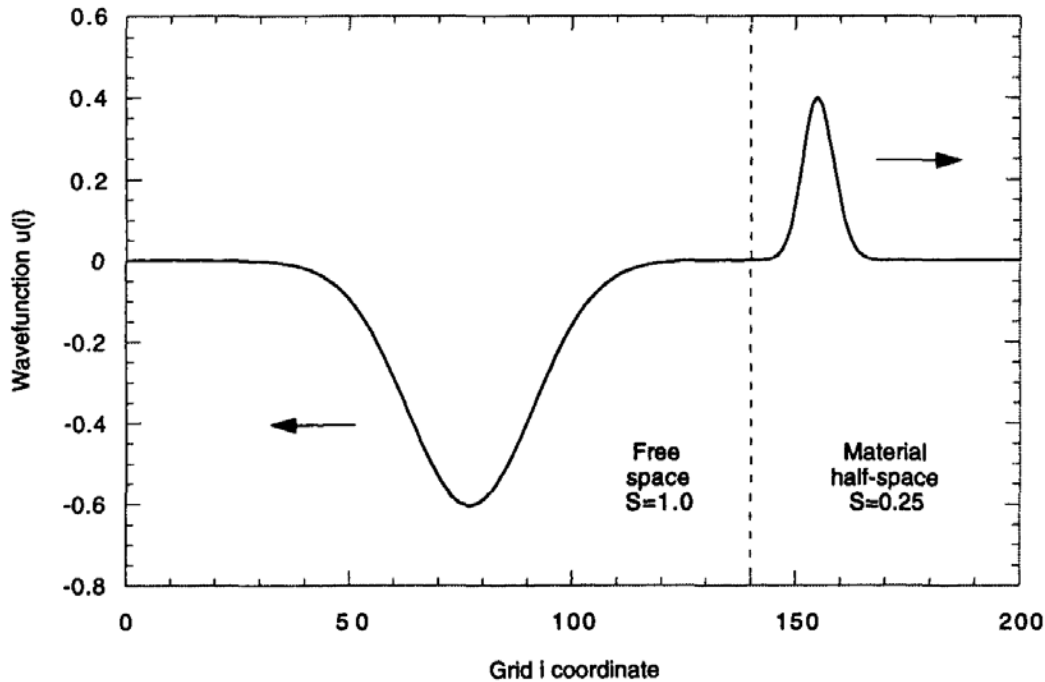
- In this case, the input signal is a Gaussian pulse which is 40 spatial cells wide at the  $1/e = 0.368$  points. From Fourier transform theory, we know that this signal has infinite frequency content, e.g.,

$$e^{-\left(\frac{t^2}{2\tau_p^2}\right)} \leftrightarrow \sqrt{2\pi} \tau_p e^{-\left(\frac{\omega^2 \tau_p^2}{2}\right)} . \text{ However, it falls off exponentially.}$$

- For the magic time-step ( $S = 1 = c\Delta t / \Delta x$ ), the pulse is perfectly recreated since  $\tilde{v}_p \equiv c$  is constant at all frequencies (no dispersion).
- When  $S = 0.99$ ,  $\tilde{v}_p$  is NOT constant at all frequencies (dispersion possible). However, since this signal has much less high frequency content, we see no visible difference with the  $S = 1$  case.

**Fig. 2.4b) comments:**

- Here, we see what happens to the Gaussian pulse when  $S = 0.5$ .
- Still no visible ringing with  $S = 0.5$ .
- Some dispersion visible in that the  $S = 0.5$  pulse lags the  $S = 1$  pulse slightly (i.e.,  $\tilde{v}_p < c$ )
- No visible superluminal signal content with  $S = 0.5$ .



**Fig. 2.5** Example of the calculated reflection and transmission of a Gaussian pulse at an interface between free space and a lossless material half-space having  $v_p = c/4$ .

**Fig. 2.5 comments:**

- In this case, the input signal was a Gaussian pulse 40 spatial cells wide at the  $1/e = 0.368$  points that is incident from free space on a lossless material where  $v_p = c/4$  at spatial index  $i = 140$ . The pulse is partially reflected ( $\Gamma = -0.6$ ) and partially transmitted ( $\tau = 0.4$ ). The grid  $\Delta x$  and  $\Delta t$  are held constant, causing the free space  $S = 1$  to change to  $S = 0.25$  in the other material.
- As can be seen, the results are very good (e.g.,  $\Gamma \approx -0.603$ , pulse width is approximately 10, a fourth of the original, in the new material, ...)

## 2.7 Numerical Stability

- Want to ensure that a bounded input results in bounded results.
- Numerical solution is unstable if a bounded input yields unbounded results.

### 2.7.1 Complex-Frequency Analysis

- assume a forward traveling sinusoidal wave in a finite-difference grid ( $\Delta x$  &  $\Delta t$ )  
 (2.6)  $u(x,t) = e^{j(\omega t - kx)}$  where  $k \rightarrow \tilde{k}$  and  $\omega \rightarrow \tilde{\omega}$ .

- In standard notation

$$\begin{aligned}
 u_i^n &= e^{j(\tilde{\omega} n \Delta t - \tilde{k} i \Delta x)} = e^{j[(\tilde{\omega}_{\text{real}} + j\tilde{\omega}_{\text{imag}})n\Delta t - \tilde{k}i\Delta x]} \\
 &= e^{-\tilde{\omega}_{\text{imag}} n \Delta t} e^{j(\tilde{\omega}_{\text{real}} n \Delta t - \tilde{k} i \Delta x)} \quad (2.40)
 \end{aligned}$$

- If  $\tilde{\omega}_{\text{imag}} > 0$ , wave decays exponentially w/ time.

- If  $\tilde{\omega}_{\text{imag}} < 0$ , the wave grows exponentially w/ time.

## 2.7.1 cont.

→ Going back to the numerical dispersion relation of (2.22), we have

$$\cos(\tilde{\omega} \Delta t) = \left(\frac{c \Delta t}{\Delta x}\right)^2 [\cos(\tilde{k} \Delta x) - 1] + 1 \quad (2.41)$$

Substituting in  $S = \frac{c \Delta t}{\Delta x}$  (Courant stability factor) and solving for  $\tilde{\omega}$ , yields

$$\begin{aligned} \tilde{\omega} &= \frac{1}{\Delta t} \cos^{-1} \left\{ S^2 [\cos(\tilde{k} \Delta x) - 1] + 1 \right\} \\ &= \frac{1}{\Delta t} \cos^{-1}(\beta) = \frac{1}{\Delta t} \left[ \frac{\pi}{2} - \sin^{-1}(\beta) \right] \quad (2.42a) \end{aligned}$$

where

$$\beta = S^2 [\cos(\tilde{k} \Delta x) - 1] + 1 \quad (2.42b)$$

→ If  $\tilde{k}$  is real,  $-1 - 2S^2 \leq \beta \leq 1$ .

(no attenuation  
due to  $\tilde{k}$ )

a) when  $0 \leq S \leq 1$ ,  $-1 \leq \beta \leq 1$  and  $\tilde{\omega}$  is real → STABLE.

b) when  $S > 1$ ,  $-1 - 2S^2 \leq \beta < -1$  →  $\tilde{\omega}$  is complex since  $\sin^{-1}(\beta)$  will be complex.

b) cont.

\* The most negative  $\beta$  occurs when the  $\cos(\tilde{\kappa}\Delta x)$  term of (2.42b) is  $-1$ , i.e.  $\tilde{\kappa}\Delta x = \pi$  or  $\tilde{\lambda} = 2\Delta x$ . So,

$$\beta_{\text{lower bound}} = 1 - 2S^2 \text{ for } \tilde{\kappa}\Delta x = \pi$$

\* Also, in this case,

$$\sin^{-1}(\beta) = -j \ln(j\beta + \sqrt{1-\beta^2}) \quad (2.44)$$

Using (2.44) in (2.42a) gives

$$\begin{aligned} \tilde{\omega} &= \frac{1}{\Delta t} \left[ \frac{\pi}{2} + j \ln(j\beta + \sqrt{1-\beta^2}) \right] \\ &= \frac{1}{\Delta t} \left[ \frac{\pi}{2} + j \ln(j\beta + j\sqrt{\beta^2-1}) \right] \quad (2.45a) \end{aligned}$$

Using  $j = e^{j\pi/2}$ , we can show that

$$\tilde{\omega} = \frac{1}{\Delta t} \left\{ \pi + j \ln(-\beta - \sqrt{\beta^2-1}) \right\} \quad (2.45c)$$

$$= \tilde{\omega}_{\text{real}} + j \tilde{\omega}_{\text{imag}}$$

2.7.1 cont. case b) cont.

Comparing terms

$$\tilde{\omega}_{real} = \frac{\pi}{\Delta t} \quad \text{and} \quad \tilde{\omega}_{imag} = \frac{1}{\Delta t} \ln(-\beta - \sqrt{\beta^2 - 1}) \quad (2.46)$$

Putting this back into our  $u_i^n$  expression yields

$$u_i^n = e^{-n \ln(-\beta - \sqrt{\beta^2 - 1})} e^{j \left[ \frac{\pi}{\Delta t} (n \Delta t) - \tilde{\kappa} i A X \right]}$$

$$u_i^n = \left( \frac{1}{-\beta - \sqrt{\beta^2 - 1}} \right)^n e^{j \left[ \frac{\pi}{\Delta t} (n \Delta t) - \tilde{\kappa} i 0 X \right]} \quad (2.47)$$

Define a multiplicative factor  $g_{\text{growth}}$

$$g_{\text{growth}} \equiv \frac{1}{-\beta - \sqrt{\beta^2 - 1}} = -\beta + \sqrt{\beta^2 - 1} \quad (2.48)$$

Remember  $\beta < -1$  for case b), so

$g_{\text{growth}} > 1 \Rightarrow$  exponential growth w/  $n$ !

Case b) cont.

Looking at the worst case  $\beta_{\text{lower bound}} = 1 - 2S^2$  where  
 $\tilde{\lambda} \Delta x = \pi$  or  $\tilde{\lambda} = 2\Delta x$ , we get

$$\begin{aligned} \rho_{\text{growth}}^{\text{max}} &= -(1 - 2S^2) + \sqrt{(1 - 2S^2)^2 - 1} \\ &= (S + \sqrt{S^2 - 1})^2 \end{aligned} \quad (2.49)$$

→ If  $S = 1$ ,  $(\rho_{\text{growth}})_{\text{max}} = 0$

→ If  $S = 1.0005$ ,  $\rho_{\text{growth}} = 1.0653$

$$(1.0653)^{10} = 1.8822$$

$$(1.0653)^{100} = 558.7$$

⋮

↓ explodes!

→ The corresponding frequency  $f_0$  to  $\tilde{\omega}_{\text{real}} = \frac{\pi}{\Delta t}$

$$\text{is } f_0 = \frac{\tilde{\omega}_{\text{real}}}{2\pi} = \frac{\pi/\Delta t}{2\pi} = \frac{1}{2\Delta t} \leftarrow \text{freq of instability!} \quad (2.50)$$

2.7.1 cont.

33

case b) cont.

→ corresponding numerical wave velocity is

$$\tilde{v}_p = \frac{\tilde{\omega}_{\text{real}}}{k} = \frac{(\pi/\Delta t)}{(\pi/\Delta x)} = \frac{\Delta x}{\Delta t} = \frac{c}{S} \quad (2.51)$$

Since  $S > 1$  in this case,  $\tilde{v}_p < c$  for the instability.

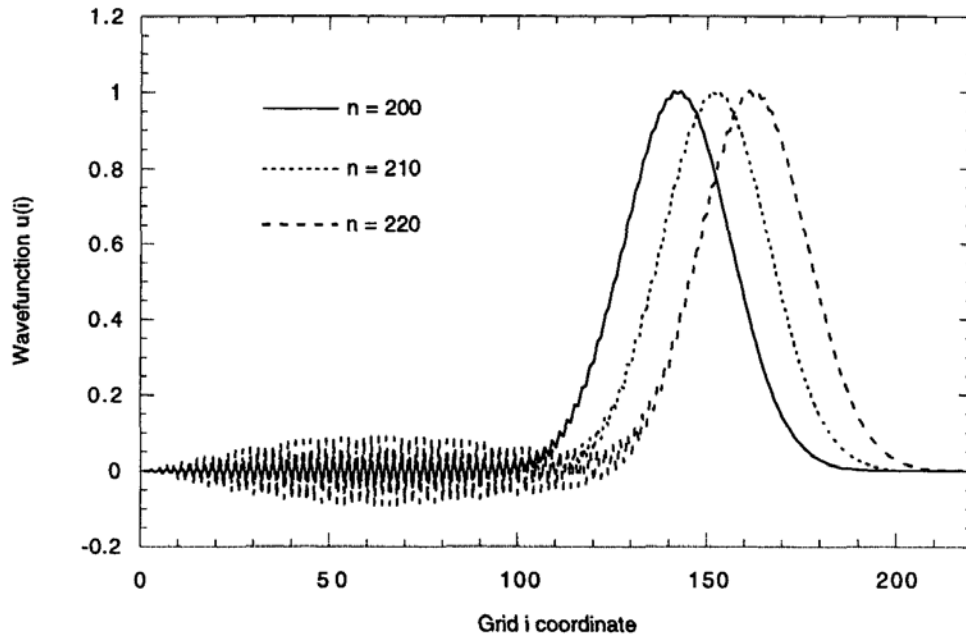
For stability,  $S \leq 1$  or  $\Delta t \leq \frac{\Delta x}{c}$  (2.52)

Usually, choose  $\Delta x + S$  then

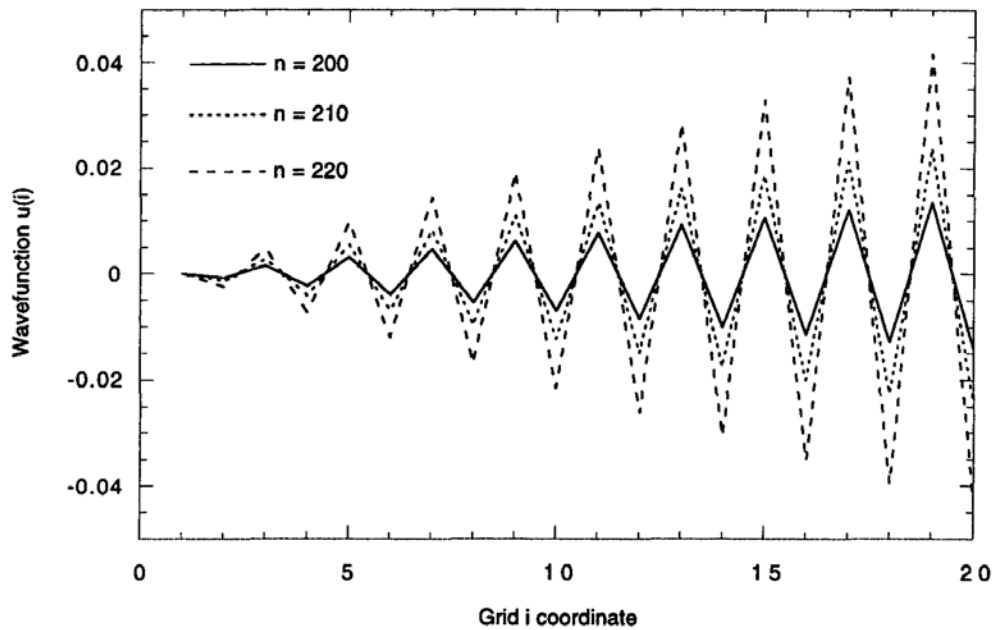
$$\Delta t = \frac{S \Delta x}{c} \quad \text{or} \quad \Delta t \leq \frac{\Delta x}{c} \quad \text{for} \\ \text{Stability}$$

\* Careful choosing  $S=1$ , any numerical round-off upward results in instability.

## 2.7.2 Examples of Calculations Involving Numerical Instability



(a) Comparison of calculated pulse propagation at  $n = 200$ ,  $210$ , and  $220$  time-steps over grid coordinates  $i = 1$  through  $i = 220$ .



(b) Expanded view of (a) over grid coordinates  $i = 1$  through  $i = 20$ .

**Fig. 2.6** The beginning of numerical instability for a Gaussian pulse propagating in free space. The Courant stability factor is  $S = 1.0005$  at each gridpoint.

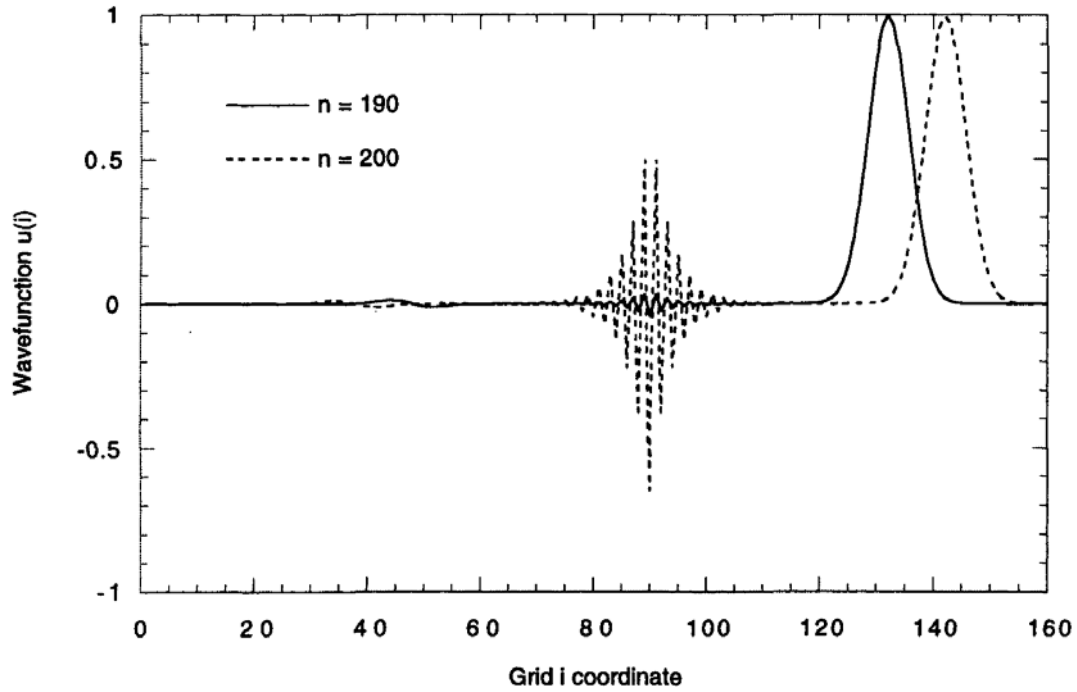
### 2.7.2 cont.

#### Fig. 2.6a) comments:

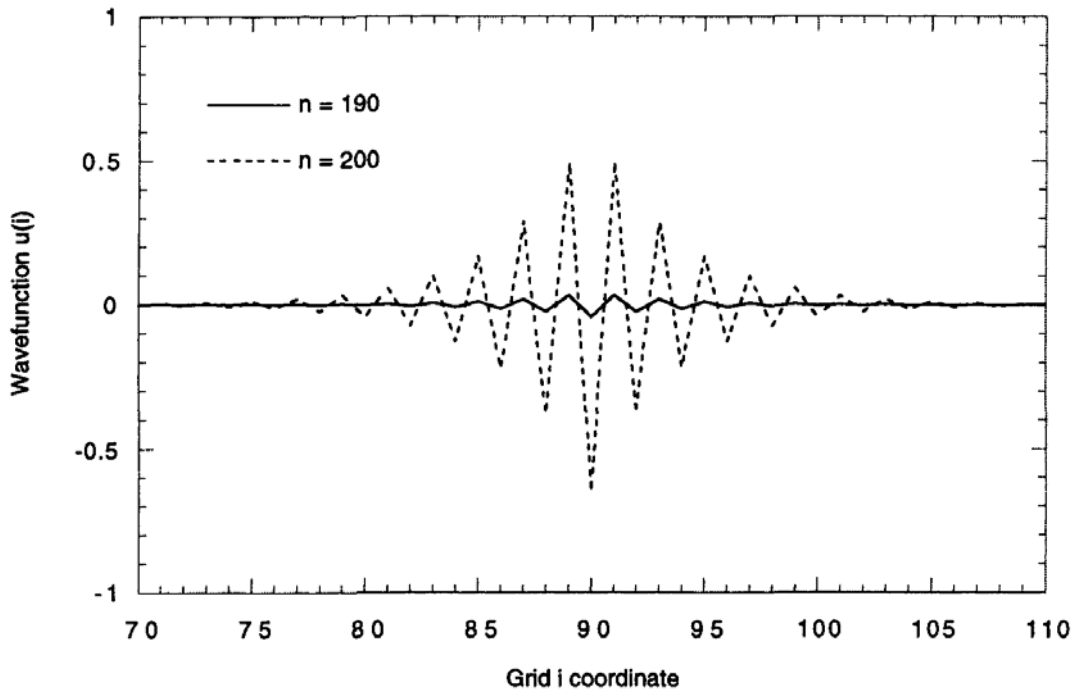
- Let  $S = 1.0005$  at all points in the grid for a Gaussian pulse input signal which is 40 temporal cells wide at the  $1/e = 0.368$  points.
- Look at several snapshots in time and see instability growing on the trailing edge of the signal.

#### Fig. 2.6b) comments:

- Here, we zoom in on the small  $i$  signal components.
- Exponential growth with respect to  $n$  is evident and near the predicted rate.
- Wavelength of instability is  $2\Delta x$  as predicted.



(a) Comparison of calculated pulse propagation at  $n = 190$  and  $n = 200$  time-steps over grid coordinates  $i = 1$  through  $i = 160$ .



(b) Expanded view of (a) over grid coordinates  $i = 70$  through  $i = 110$ .

**Fig. 2.7** The beginning of numerical instability for a Gaussian pulse propagating in free space. Unlike Fig. 2.6, the Courant stability factor is  $S = 1$  at all gridpoints but  $i = 90$ , where  $S = 1.075$ .

**Fig. 2.7a) comments:**

- Let  $S = 1.075$  at a single point ( $i = 90$ ) in the grid for a Gaussian pulse input signal which is 40 temporal cells wide at the  $1/e = 0.368$  points.  $S = 1$  elsewhere in the grid.
- Look at several snapshots in time and see instability growing, centered on  $i = 90$ , at the trailing edge of the signal.

**Fig. 2.7b) comments:**

- Here, we zoom in on the signal components near  $i = 90$ .
- Exponential growth with respect to  $n$  is evident, but is slower than predicted rate for global grid.
- Wavelength of instability is  $2\Delta x$  as predicted.